

## Minerals on Titan: the crystal structure of diacetylene

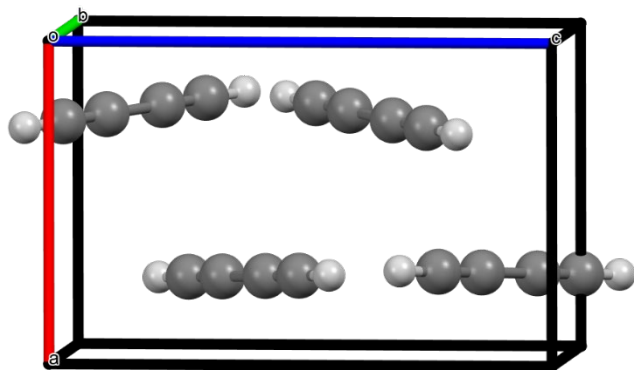
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Saturn's largest moon, Titan, is an important extra-terrestrial target for studying planetary chemistry and potential astrobiology. With an upper atmosphere rich in nitrogen and methane that undergoes conversion into hydrocarbons, nitriles and aerosols, this icy moon is a valuable model for investigating prebiotic chemistry analogous to early-Earth [1,2]. Among the larger molecules detected in Titan's atmosphere, diacetylene ( $C_4H_2$ ) is of particular interest due to its significant absorption of photons at longer wavelengths, which facilitates the formation of radical species such as  $C_4H$  and  $C_4H_3$  that are key to polycyclic aromatic hydrocarbon (PAHs) formation [3].

In this study, we combined powder X-ray and neutron diffraction, Raman spectroscopy and periodic-DFT calculations to characterise the crystal structure of diacetylene for the first time. The structure is described in the space group  $P2_12_12_1$  with four molecules in the unit cell (Fig. 1). The diacetylene molecules are arranged in a layered structure dominated by  $CH\cdots\pi$  interactions, which leads to anisotropic thermal expansion behaviour. At ambient pressure, no structural phase transitions were observed in the 5–220 K temperature range. Due to its structural similarity with acetylene, diacetylene may serve as a potential co-crystal component of particular importance for Titan's surface chemistry and geomorphology. This is of potential importance in view of the upcoming NASA Dragonfly mission.



**Figure 1.** Representation of the structure of diacetylene refined in space group  $P2_12_12_1$  against neutron powder diffraction data collected at 5 K.

[1] He, C., Smith, M. A. (2014). *Icarus*, **238**, 86-92.

[2] Raulin, F., Brassé, C.; Poch, O., Coll, P. (2012). *Chem. Soc. Rev.*, **41**, 5380.

[3] Huang, C., Zhang, F., Kaiser, R. I., Kislov, V. V., Mebel, A. M.; Silva, R., Gichuhi, W. K.; Suits, A. G. (2010). *ApJ*, **714**, 1249-1255.

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